Contributions to Forest Fire Simulations: Mathematical Models, Numerical Methods and GIS Integration

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INTRODUCTION

- Propagation models
- Combustion models
- 2 SIMPLIFIED PHYSICAL MODEL
 - The goal
 - A Simple Physical Model
 - Enthalpy operator
- 3 NUMERICAL METHOD
 - Time Integration
 - Solution at each time step
- Non Local Radiation Model
 Characteristic Method
- 5 Wind Model
- 6 Simulations
 - wind model example
 - Physical and numerical data



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Clasification

Models

PROPAGATIONCOMBUSTIONPosition of the fire frontModelization of physics

models with increasing complexity:

- Cellular Automata
- Geometric
- Empiric models
- Reaction Diffusion Convection



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Cellular Automata

- Cell with several states (burning, burnt, not burning)
- Transition probabilities as a function of the neighbours cells
 - Advantages: \implies Fast computation
 - Disadvantages: It modelize probabilistic phenomena Not direct relation with physical parameters





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Geometric models

- 1D Fire front on a 2D surface
- Huygens principle
- Advantages and disadvantages the same as automata.





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Empiric models

- An energy balance is considered on the fire front
- Rate of spread given by empiric laws
- R.C.Rothermel, *A Mathematical Model for Predicting Fire Spread in Wildland fuels.*
- Advantages \implies Fast computation
- Disadvantages \implies Parameters must be adjusted case by case.



Combustion models: Convection-Difussion-Radiation-Reaction

Complex models

- Several phases and conservation laws are considered
- Solid phase and gas phase with different temperature
- Two layer models
- Three dimensional equations
- Large time computation

Simplified models

- An average medium is considered
- Only one temperature
- One phase (other phases parametrized)
- One or two dimensional equations considered
- Could allow real time computation or faster

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Propagation models Combustion models

Combustion models: A general example. Physical Assumptions

Porous medium of porosity ϕ and density ho

- Composed of solid fuel Y_s, oxygen Y_o and gaseous fuel Y_g
- Gases are all assumed to have an equal temperature T_g . The solid temperature T_s is coupled by the term $h(T_s T_g)$
- Solid fuel Y_s transforms into gaseous fuel Y_g through pyrolysis.
- 5 Gaseous fuel Y_g reacts with the oxygen Y_o
- 6 Which generates the flames and heat
- Gases and temperatures are under the influence of convection, diffusion and heat is loss in the vertical direction (cooling)
- 8 Gain in solid temperature due to radiation R and combustion heat (proportional to the solid fuel combusted)



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- $(1-\phi)\rho_s C_s \frac{\partial T_s}{\partial t} \nabla(\kappa_s \nabla T_s) = R(T_g) h(T_s T_g)$

- T_s Temperature of the solid



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- T_g Temperature of the gases
- Y_s Solid fuel
- Y_o Oxygen
- Y_g Gaseous fuel



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Time terms

- Convective transport due to the wind
- Diffusion within the vegetation
- Vertical cooling or vertical loss
- Link between T_s and T_g
- Arrhenius type laws for combustion and pyrolysis
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$$\begin{array}{l} \bullet (1-\phi)\rho_s C_s \frac{\partial T_s}{\partial t} - \nabla(\kappa_s \nabla T_s) = R(T_g) - h(T_s - T_g) \\ \bullet \rho C(\frac{\partial T_g}{\partial t} + \mathbf{v} \cdot \nabla T_g) - \nabla(\kappa_g \nabla T_g) = -h(T_g - T_s) + qA_g Y_o Y_g e^{-\frac{E_g}{RT_g}} - h_s(T_g - T_\infty) \\ \bullet \frac{\partial Y_s}{\partial t} = -A_s Y_s e^{-\frac{E_g}{RT_s}} \\ \bullet \frac{\partial Y_o}{\partial t} + \mathbf{v} \cdot \nabla Y_o - \nabla(\kappa_{oo} \nabla Y_o) = -h_o(Y_o - Y_{o,\infty}) - A_g Y_o Y_g e^{-\frac{E_g}{RT_g}} \\ \bullet \frac{\partial Y_g}{\partial t} + \mathbf{v} \cdot \nabla Y_g - \nabla(\kappa_{gg} \nabla T_s) = -h_g(Y_g - 0) - A_g Y_o Y_g e^{-\frac{E_g}{RT_g}} + AY_s e^{-\frac{E_g}{RT_s}} \end{aligned}$$

- Time terms
- Convective transport due to the wind
- Diffusion within the vegetation
- Vertical cooling or vertical loss
- Link between T_s and T_g
- Arrhenius type laws for combustion and pyrolysis
- Radiation term $R(T_g)$



1 INTRODUCTION

- Propagation models
- Combustion models

2 SIMPLIFIED PHYSICAL MODEL

- The goal
- A Simple Physical Model
- Enthalpy operator

3 NUMERICAL METHOD

- Time Integration
- Solution at each time step
- Non Local Radiation Model
 Characteristic Method
- 5 Wind Model
- 6 Simulations
 - wind model example
 - Physical and numerical data



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Numerical Simulation of forest fires in small computers with computation times being a small fraction of the real time





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Simple models

- Simplified mathematical models
- Basically two dimensional models
- Some physical phenomena will be parametrized.

Realistic models

- Take into account main mechanisms of propagation
- Take into account three dimensional effects

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Convection-Diffusion-Radiation simplified models

Require the solution 1D or 2D of Convection-Diffusion-Radiation equations.

layer conducting radiating material

•
$$\rho C(\frac{\partial u}{\partial t} + \mathbf{V} \nabla u) - (J * u^4 - u^4) + \alpha u = Af(u, y)$$

•
$$\frac{\partial y}{\partial t} = -f(u, y)$$

- *u* Temperature
- y Mass fraction of fuel
- $-(J * u^4 u^4)$ Non local diffusion (due to radiation)



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Local approximation

Estimation of the term
$$-(J * u^4 - u^4)$$

$$(J * u^{4})(x) - u^{4} = \int J(x - y)u^{4}(y)dy - u^{4}(x)$$

= $\int J(x - y)(u^{4}(y) - u^{4}(x))dy$
= $\int J(x - y)((u^{4})'(x)(y - x) + \frac{1}{2}(u^{4})''(x) + ...)dy$
= $-(\int J(z)zdz)(u^{4})'(x) + (\int J(z)z^{2}dz)(u^{4})''(x) + ...)dy$

If the kernel J(.) is symmetric then

$$-(J * u^4 - u^4) \approx -(\int J(z)z^2 dz)(4u^3u')'(x)$$



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If the kernel is not symmetric (due to wind which tilt the flame) a convection term appears

$$-(J * u^{4} - u^{4}) \approx (\int J(z)zdz)(u^{4})'(x) - (\int J(z)z^{2}dz)(4u^{3}u')'(x)$$



Diffusion approximation

If it is assumed that the absortion mean penetration distance $\frac{1}{a}$ is small compared with the distances over which significant temperatures changes occur, that is the effect of radiation is local (Rosseland approximation)

layer conducting radiating material

•
$$\rho C(\frac{\partial u}{\partial t} + \mathbf{V}\nabla u) - \nabla(\frac{16\sigma u^3}{3a_R}\nabla u) + \alpha u = Af(u, y)$$

• $\frac{\partial y}{\partial t} = -f(u, y)$



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A Simple Physical Model

The non dimensional equations governing the fire spread in a region $\boldsymbol{\Omega}$ are,

- 2 $e \in G(u)$
- $\partial_t y = -g(u)y$
- r(u, y) is the source term
- G(u) is the Enthalpy
- g(u)y is the rate of pyrolysis



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$$G(u) = \begin{cases} u & \text{if } u < u_{v} \\ [u_{v}, u_{v} + \lambda_{v}] & \text{if } u = u_{v} \\ u + \lambda_{v} & \text{if } u_{v} < u < u_{p} \\ [u_{p} + \lambda_{v}, u_{p} + \lambda_{v} + \lambda_{p}] & \text{if } u = u_{p} \\ u + \lambda_{v} + \lambda_{p} & \text{if } u > u_{p} \end{cases}$$





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Comments

- Local diffusion has been neglected
- Convection has been neglected but main WIND EFFECTS are considered via radiation model
- Main mechanism of fuel heating is **RADIATION** (non local diffusion)
- Effects of the WATER CONTENT are considered
- Gas phase is parametrized (The temperature and height of the flame are parameters)



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 - Time Integration
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Euler semi-implicit method

Let $\Delta t = t^{n+1} - t^n$ a time step and let y^n , e^n and u^n denote approximations at time step t^n , to the exact solution y, e and u respectively.

We consider a semi-implicit scheme. At each time step we solve,

$$\frac{e^{n+1}-e^n}{\Delta t} + \alpha u^{n+1} = r^n$$
$$e^{n+1} \in G(u^{n+1})$$
$$\frac{y^{n+1}-y^n}{\Delta t} = -y^{n+1}g(u^{n+1})$$



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Solving the nonlinear equation in each time step

The multivalued operator G(.) is maximal monotone, then its resolvent $J_{\lambda} = (Id + \lambda G)^{-1}$ for any $\lambda > 0$ is a well defined univalued operator.

Taking $\lambda = 1/(\alpha \Delta t)$ the algorithm simplifies

$$u^{n+1} = J_{1/\alpha\Delta t} \left(\frac{1}{\alpha\Delta t}e^n + \frac{1}{\alpha}r^n\right)$$

$$e^{n+1} = e^n - \alpha\Delta t u^{n+1} + \Delta t r^n$$

$$y^{n+1} = \frac{y^n}{1 + \Delta t r(u^{n+1})}$$



Practical Computation of the Resolvent

 $(\alpha \Delta t \ I + G) \overline{z} \ni \overline{u}$





$$\bar{z} = u_{v}$$

• if $(1 + \alpha \Delta t) u_{v} + \lambda_{v} < \bar{u} < (1 + \alpha \Delta t) u_{p} + \lambda_{v}$
$$\bar{z} = \frac{\bar{u} - \lambda_{v}}{1 + \alpha \Delta t}$$

Figure: Computation of \overline{z}



Nonlocal radiation

The source term is due to nonlocal radiation: In each point it is caculated summing up the radiation intensity for all directions Ω .



1- Assume as known $i(x, \Omega) \quad \forall \Omega$ for a given x

2- By use of polar coordinates at the tangent space of $\bar{\mathbf{x}}$ and defining $\mu = \cos \theta$ and $\gamma = \cos \phi$

 $r(ar{\mathbf{x}}) = \int i(ar{\mathbf{x}}, \mathbf{\Omega}) \mathbf{\Omega}. \mathbf{n} \; d\omega = \int_{\mu, \gamma} rac{i_+(ar{\mathbf{x}}, \mu, \gamma)\mu}{\sqrt{1-\gamma^2}} \; d\mu d\gamma$

3- A quadrature may be done with respect to:

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ightarrow ext{Gauss-Chevichev}$



Radiation equation

After adimensionalization, the radiation equations in the direction Ω can be written as

 $\mathbf{\Omega} \cdot \nabla i + \mathbf{a}^* i = \delta (1 + u_g)^4 \quad \text{in } D$ $i = 0 \quad \text{on } \partial D \cap \{\mathbf{x}; \ \mathbf{\Omega} \cdot \mathbf{n} < 0\}$

The incident energy at a point $\mathbf{x}(x, y, h(x, y))$ of the surface S per unit time and per unit area will be obtained summing up the contribution of all directions Ω

$$r(\mathbf{x}) = \int_{\omega=0}^{2\pi} i(\mathbf{x}, \mathbf{\Omega}) \mathbf{\Omega}.\mathbf{n} \ d\omega \tag{1}$$

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The radiation term r in the energy equation is computed by numerical integration.

Summing up for all the solid angles

$$r(\bar{\mathbf{x}}) = \int_{\theta=0}^{\theta=\pi/2} \int_{\phi=0}^{\phi=2\pi} i(\bar{\mathbf{x}},\theta,\phi) \cos\theta \sin\theta \ d\theta d\phi = \int_{\mu=0}^{\mu=1} \int_{\gamma=-1}^{\gamma=1} \frac{i_{+}(\bar{\mathbf{x}},\mu,\gamma)\mu}{\sqrt{1-\gamma^{2}}} \ d\mu d\gamma + \int_{\mu=0}^{\mu=1} \int_{\gamma=-1}^{\gamma=1} \frac{i_{-}(\bar{\mathbf{x}},\mu,\gamma)\mu}{\sqrt{1-\gamma^{2}}} \ d\mu d\gamma \quad (2)$$

 $\mu = \cos \theta$, $\gamma = \cos \phi$ i_+ (resp. i_-) stands for the radiation intensity *i* corresponding to an angle ϕ such that $0 \le \phi < \pi$ (resp. $\pi \le \phi < 2\pi$).



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The integrals are computed using Gauss-Legendre quadrature with respect μ and Gauss-Chebyshev quadrature with respect γ in order to cope with the singular weight $\frac{1}{\sqrt{1-\gamma^2}}$.

$$r(\bar{\mathbf{x}}) \approx \sum_{k,l} W_{kl} i_{+}(\bar{\mathbf{x}}, \mu_{k}, \gamma_{l})\mu_{k} + \sum_{k,l} W_{kl} i_{-}(\bar{\mathbf{x}}, \mu_{k}, \gamma_{l})\mu_{k}$$
(3)



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Numerical solution of the radiation equation

The characteristic line is

 $[0,\xi] \longmapsto \mathcal{R}^3$

 $\xi \longrightarrow (x(\xi) = \bar{x} + \xi \Omega_1, \ y(\xi) = \bar{y} + \xi \Omega_2, \ z(\xi) = \bar{z} + \xi \Omega_3)$

On the characteristic, the radiation equation becomes

$$\frac{di}{d\xi} + a^*i = \delta(1+u_g)^4$$

which can be solved backwards together with the condition

$$\lim_{\xi\to\infty}i(\xi)=0$$

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Extension of the temperature field

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When integrating the radiation equation we need to know $a^*(\xi(x, y, z))$ and $u_g(\xi(x, y, z))$ so we need to extend the temperature field to the whole domain D.

• If there is NO WIND, we extend the temperature vertically

 $\tilde{u}(x, y, z) = u(x, y, h(x, y))$

h(x,y) < z < h(x,y) + H

• In the case of WIND CONDITIONS we compute the extended field assuming a convective transport

 $\tilde{u}(x, y, z) = u(x - (z - h(x, y))\frac{v_x}{v}, y - (z - h(x, y))\frac{v_y}{v}, h(x, y))$



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The radiation equation and complex geometrie

The method is able to cope with complex geometries





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Vertical diffusion model



Gives the horizontal wind field in a 3D domain by the expression

$$V(\mathbf{x}, z) = m(\mathbf{x}, z)\nabla p(\mathbf{x}) + n(\mathbf{x}, z)\nabla \hat{T}(\mathbf{x})$$

•
$$m(\mathbf{x}, z) = \frac{1}{2}z^2 - \delta z - \frac{1}{2}h^2(\mathbf{x}) + (\delta + \xi)h(\mathbf{x}) - \xi\delta$$

• $n(\mathbf{x}, z) = -\frac{1}{24}z^4 + \frac{1}{6}\delta z^3 - \frac{1}{3}\delta^3 z + \frac{1}{24}h^4(\mathbf{x}) - \dots$
• $n(\mathbf{x})$ is a potential function



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Vertical diffusion model

The potential $p(\mathbf{x})$ satisfies the following boundary problem

$$\begin{aligned} -\nabla(a\nabla p) &= \nabla(b\nabla\hat{T}) \quad \text{in } d\\ a\frac{\partial p}{\partial n} &= -b\frac{\partial\hat{T}}{\partial\nu} + (\delta - h)v_m.\nu \quad \text{on} \quad \partial d \end{aligned}$$

•
$$a(\mathbf{x}) = \frac{1}{3}(\delta - h(\mathbf{x}))^2(3\xi + \delta - h(\mathbf{x}))$$

• $b(\mathbf{x}) = \frac{1}{30}(\delta - h(\mathbf{x})^2(2\delta^2(2\delta + 5\xi) - 2\delta(\delta - 5\xi)h(\mathbf{x}) - (3\delta + 5\xi)h^2(\mathbf{x}) + h^3(\mathbf{x}))$



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Optimal control problem

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The optimal control problem is caracterised by

$$\int_{\omega} a \nabla p(u) \nabla \varphi + \frac{1}{\alpha} \int_{\partial \omega} q \varphi = -\int_{\omega} b \nabla \hat{T} \nabla \varphi \quad \forall \varphi \in V$$

$$\int_{\omega} a \nabla q(u) \nabla \psi$$

$$-\sum_{i=1}^{N} \int_{\omega} \rho_{\epsilon} (x - x_{i}) (m \nabla p(u) + n \nabla \hat{T} - V_{i}) m \nabla \psi = 0 \quad \forall \psi \in V$$

$$u = -\frac{1}{\alpha} q \quad \text{on} \quad \partial d$$

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wind model example Physical and numerical data

An example: Simulation: Cofrentes (Valencia, Spain)

The data

- Zone 5 Km long by 5 Km width
- Average Meteorological wind given.
- The ignition point
- GIS data: Surface orography, vegetation density
- Combustion properties of the fuel: half-life decay.



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Data by courtesy of Tecnosylva, S.L. León, Spain





Figure: Wind Module



Figure: Wind direction

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Figure: With wind at time 33 minutes

Figure: Without wind at time 33 minutes







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Figure: With wind at time 66 minutes

Figure: Without wind at time 66 minutes







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Figure: With wind at time 99 minutes

Figure: Without wind at time 99 minutes







Figure: With wind at time 132 minutes

Figure: Without wind at time 132 minutes

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Figure: With wind at time 165 minutes

Figure: Without wind at time 165 minutes

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Figure: With wind at time 198 minutes

Figure: Without wind at time 198 minutes

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Figure: With wind at time 231 minutes

Figure: Without wind at time 231 minutes

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Figure: With wind at time 264 minutes

Figure: Without wind at time 264 minutes

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Figure: With wind at time 297 minutes

Figure: Without wind at time 297 minutes

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wind model example Physical and numerical data

3D representation



Figure: Without wind



Figure: With wind

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An example: Simulation of a fire in a river basin

The data

- Zone 6 Km long by 3 Km width
- Meteorological wind given in several points.
- The ignition point
- GIS data: Surface orography, vegetation density
- Combustion properties of the fuel: half-life decay.





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Wind: wind field V

velocidad, z=0.1+h





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Initial Fuel Density





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Main physical parameters

- Zone 6 Km long by 3 Km width
- Water content: 2%
- Half-time decay: 700 seconds
- Flame height: 20 meters
- Flame temperature 1225C
- Wind velocity (1st): NO WIND
- Wind velocity (2nd): 10m/s ÷ 20m/s

Main numerical data

- Mesh size: maximum size 100 meters, minimum size 50 meters
- number of uknowns per variable: 31774
- Time step: 200 seconds
- Numerical integration points for radiation: 8 × 2 = 16

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Computation results

- Real time of the simulation \approx 5 hours 30 minutes
- Computation time in my laptop: 25 minutes
- Mean fire front velocity: 650 m/h (no wind)







Computation results

- Real time of the simulation \approx 5 hours 30 minutes
- Computation time in my laptop: 25 minutes
- Mean fire front velocity: 710 m/h (wind)





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Outline INTRODUCTION SIMPLIFIED PHYSICAL MODEL wind model example Physical and numerical data

Evolution of a fire in the Ebro basin

SHOW ANIMATIONS



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L. Ferragut, M. Asensio, S.Monedero† Fire modelling and Forest Fire Decision System